

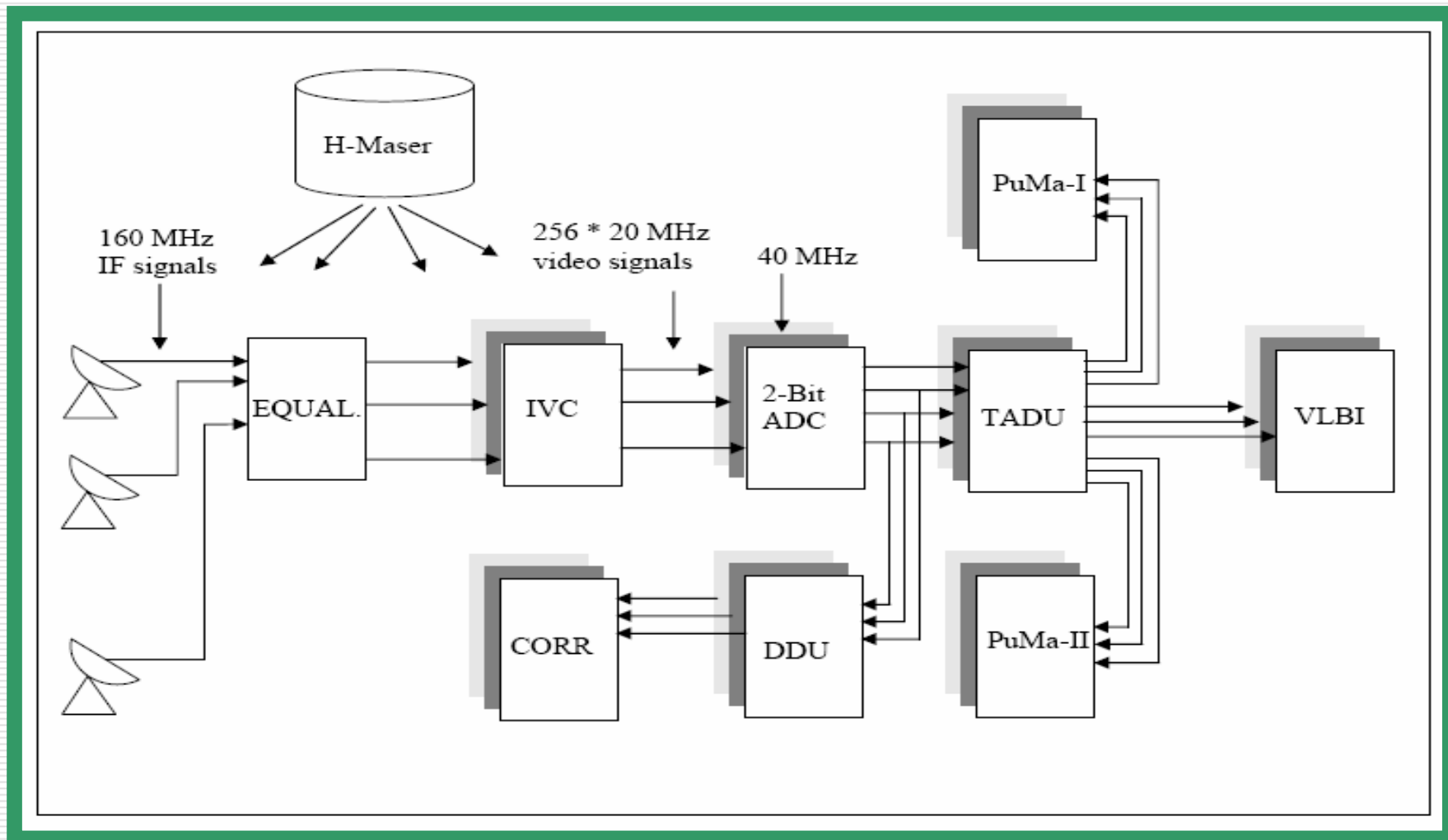
WSRT Pulsar Capabilities : features & prospects

WSRT Users Meeting,
4th June 2008,
Amersfoort.

Outline

- ❑ WSRT overview
 - ❑ WSRT Pulsar advantage
 - ❑ Pulsar Machine - II
 - ❑ PuMa-II - Current capabilities
 - ❑ Future Prospects
-

Straw man WSRT

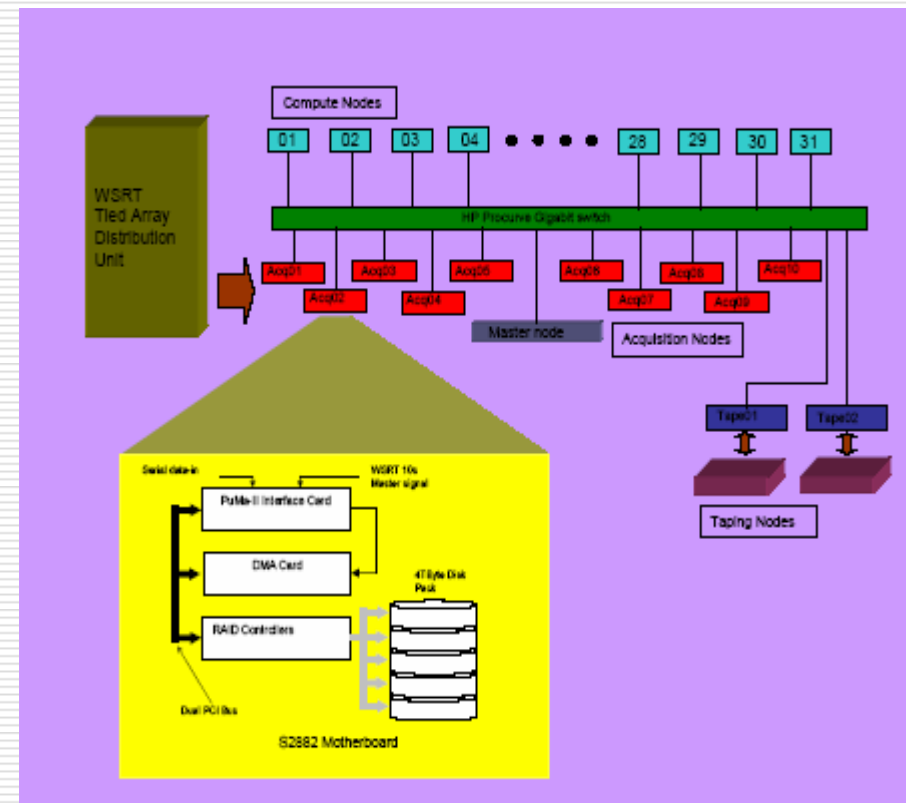


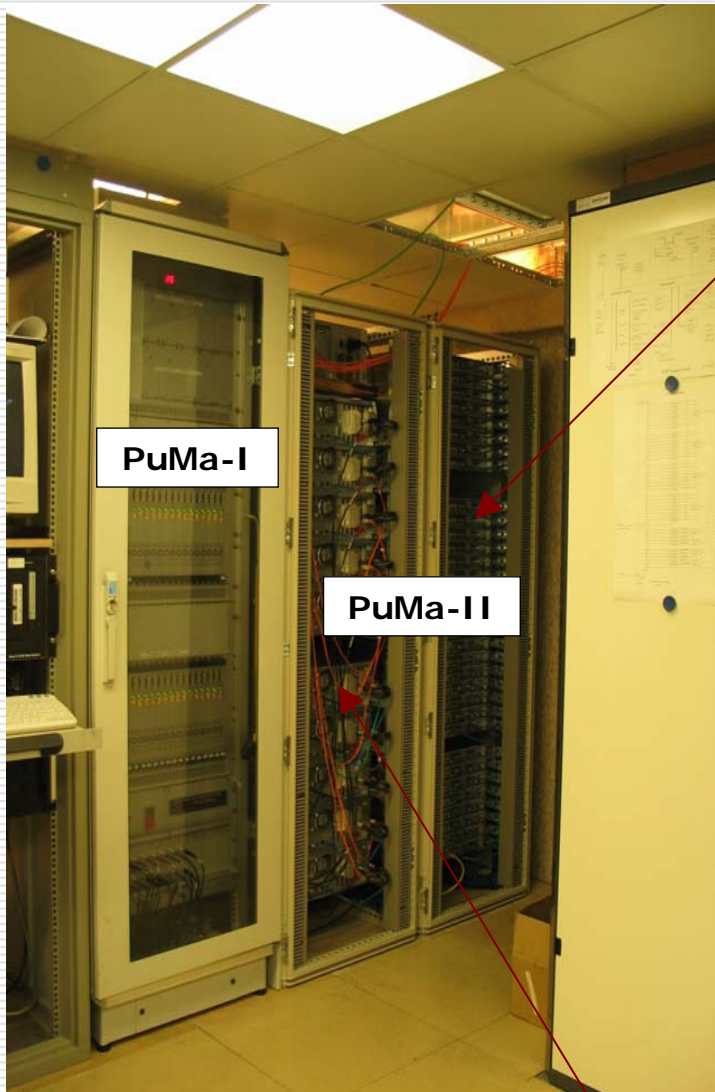
The WSRT-PuMa-II Advantage

- Sensitivity – Tied Array mode
 - Multi-Frequency (new $F_{\text{sky}} < 1\text{min.}$)
 - Multi-subarray mode
 - Pulsar Machine II
 - Flexible Beowulf-like cluster
 - 8 bit system
 - Baseband recorder $\sim 8 \times 20$ MHz
 - $\sim 50\text{ns}$ time resolution
 - New 2-bit, 160 MHz possibility
 - Replace PuMa-I (end 2008)
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PuMa-II Overview

- 44-node cluster
 - dual 2.0 GHz CPUs
 - 10 storage nodes (27TB)
 - 31 compute node
 - 2 archival nodes (22 TB)
 - ≈ 500 GFlops
 - ≈ 70 TB storage
- 640 Mbytes/sec throughput
- Optical interface to WSRT
 - Signal quality
- Accurate time stamps
 - H-Maser



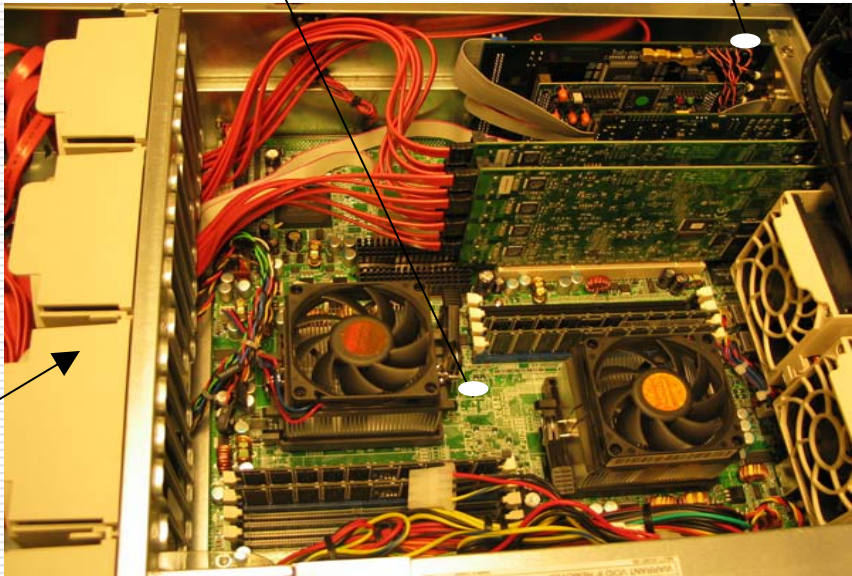


Compute nodes



Dual 2.0 GHz Opetrons
4TB RAID'ed Disks
2GB RAM
Dual PCI Bus
210 MB DMA card

Storage nodes

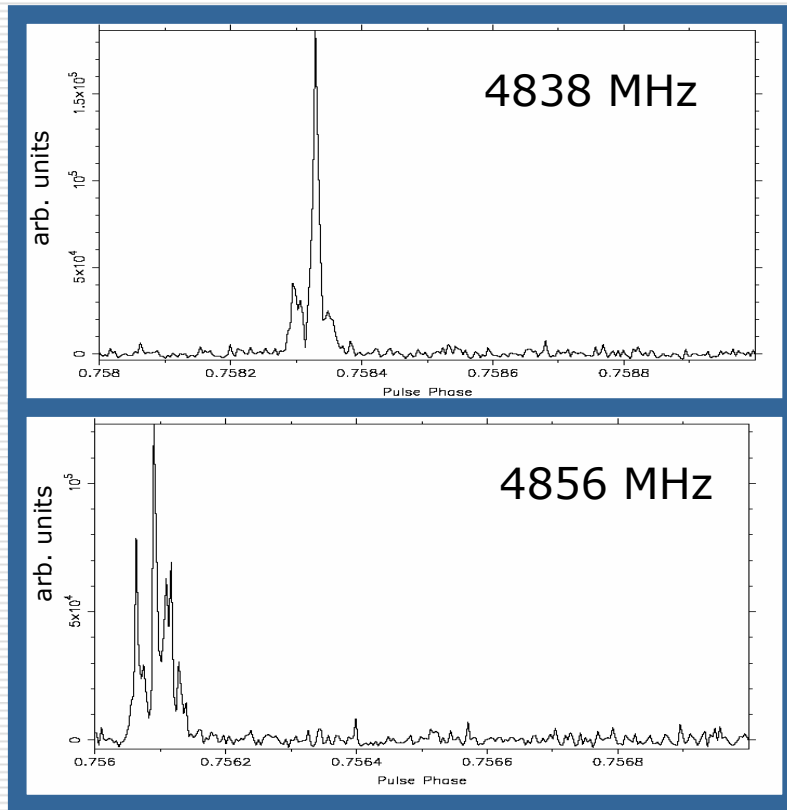


PuMa-II : Current capabilities

- Distributed data recording
 - C-based acq. s/w: 8 storage nodes – 8x20 MHz
 - 12-hour recording
 - Distributed signal processing
 - Near realtime data reduction
 - Sun Grid Engine (SGE) queue system
 - DSP for pulsars (dspsr)
 - Coherent dedispersion, fold, GP search
 - PSRCHIVE + TEMPO
 - View, clean RFI, flux/poln. Calibration, timing
-

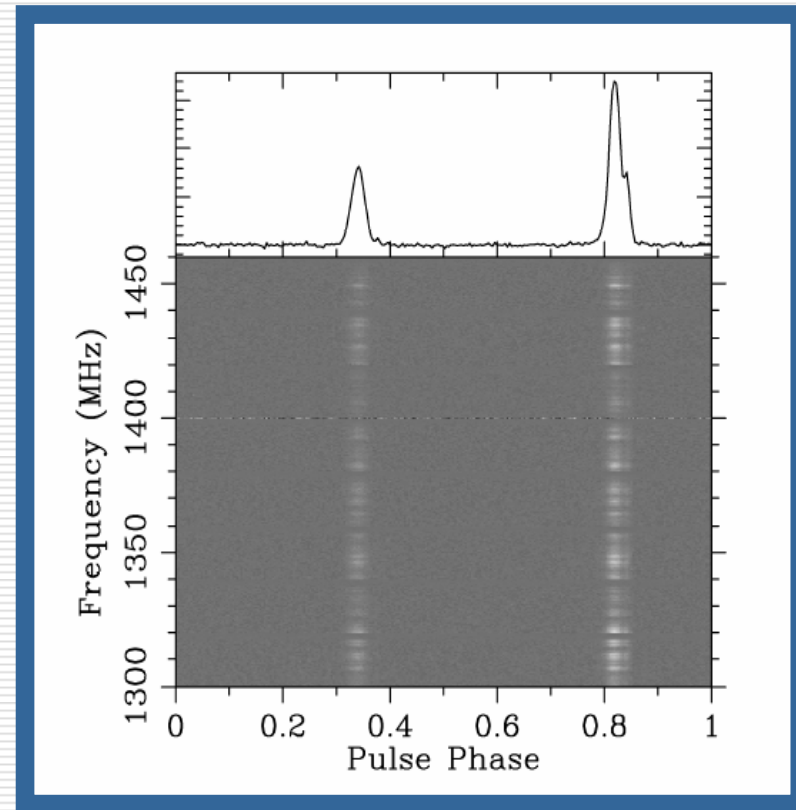
PuMa-II: Example 1

PSR B0531+21 – giant pulses



$t_{\text{res}} = 100 \text{ ns}$

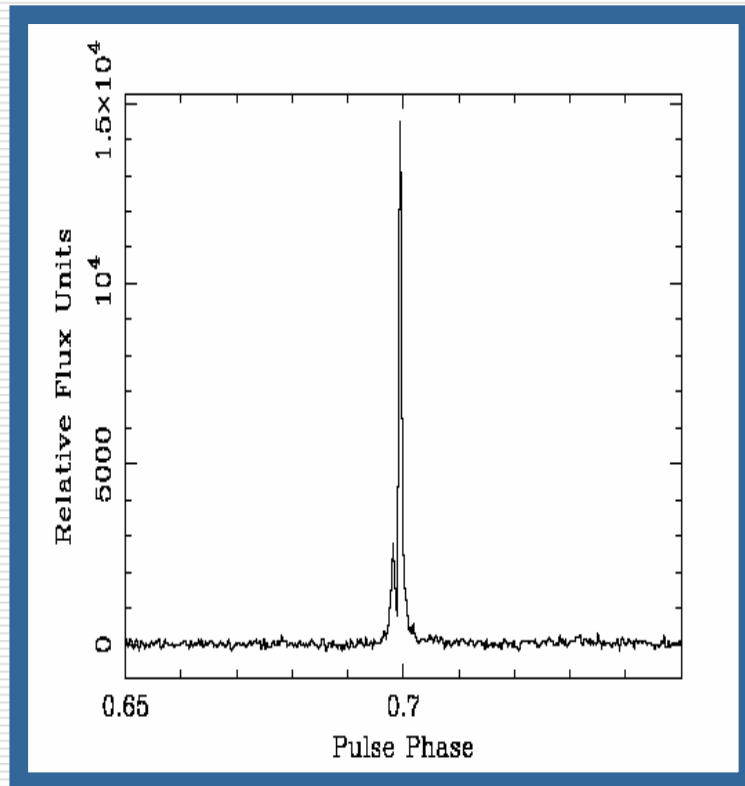
PSR B1937+21



PuMa-II TOA $\sim 60 \text{ ns}$

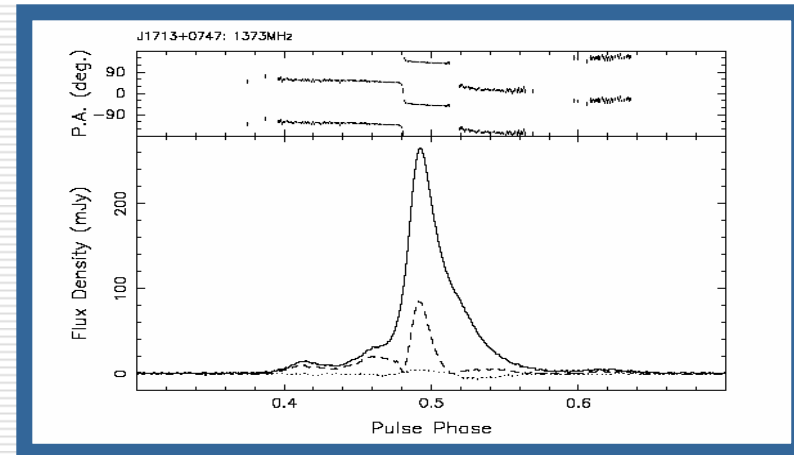
PuMa-II: Example 2

PSR B1133+16 – GPs?

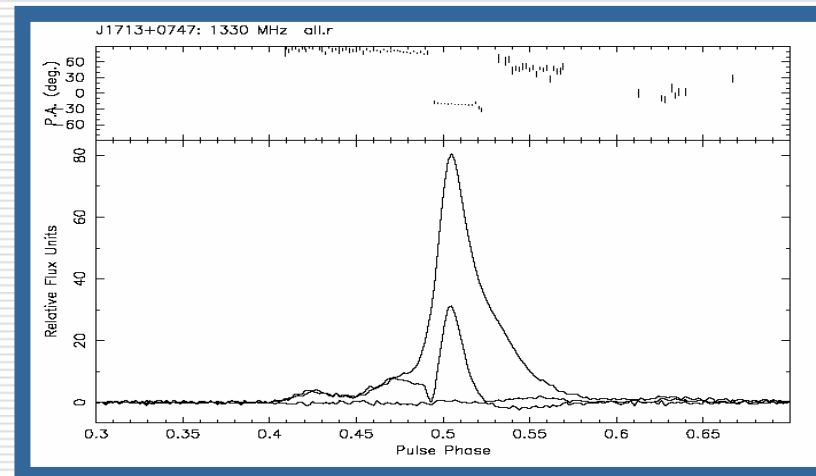


$T_{\text{res}} = 256 \mu\text{s}$, peak flux $\sim 2000 \text{ Jy}$

PSR J1713+0747
Polarization.



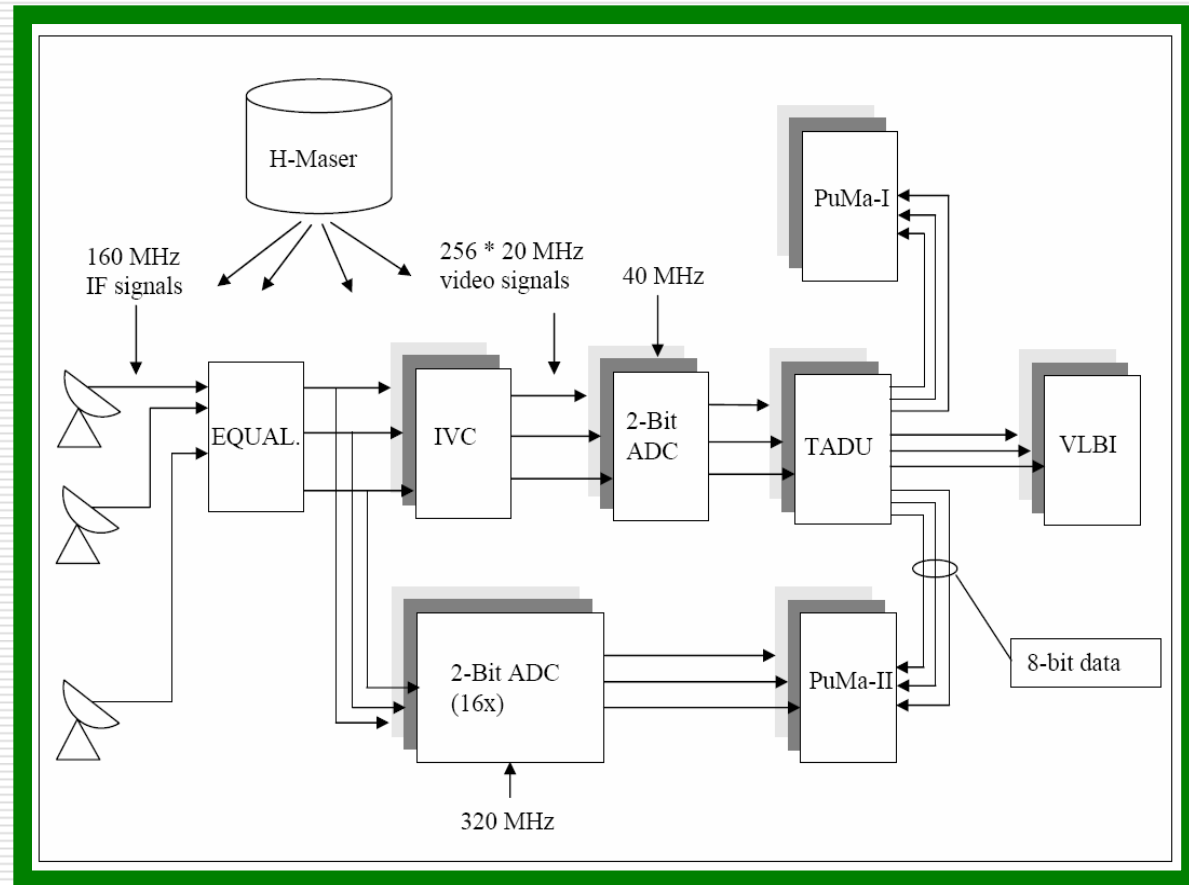
Ord et. al, MNRAS, 2004



$\sim 10\%$ more linearly poln. flux

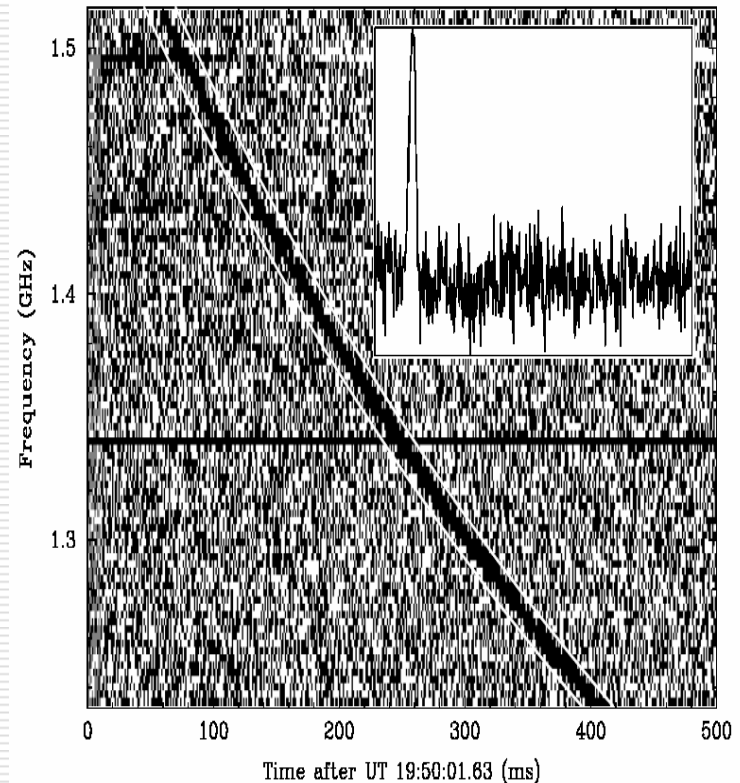
PuMa-II: Future prospects

- 2 Bit Mode
 - $\sim 3\text{ns}$
 - 320 MHz clock
 - Reuse software
- Incoh. Addition
 - $S/N \sim \sqrt{14}$
- Transient monitor
 - Large FoV
- Software Corr.
- APERTIF signal processing test bed



2Bit Mode : Science

- Eg. ~5ms “Lorimer burst”
 - 30Jy, 1.5 Gpc (DM=375)
 - Probe Intergalactic medium
- Piggy-back monitor mode
 - Better use of telescope
- Reduce location uncertainty
 - Software correlator
- LOFAR trigger
 - eg: DM=500@1400MHz → 40 min. delay @ 40 MHz



Lorimer et. al, arXiv:0709.4301 (2007)