2025-09-09 01:50 1/7 makesourcedb

# makesourcedb

A source table is used by BBS to know the available sources and their parameters. It is a special form of a Parameter Table and is described in detail in parmdb.pdf .

The sources in a source table can optionally be grouped into so-called patches. BBS makes use of patches by, for instance, combining them when doing parameter fits. It is possible in makesourcedb to define patches and tell to which patch a source belongs. If no patch is defined for a source, the source is its own patch. Thus for each such source a patch is created, where it is possible to add a prefix and/or suffix to the patch name making it possible to use patch patterns in BBS.

The program makesourcedb can be used to fill or append a source table with (selective) sources read from an ASCII file. Note that the script gsm.py can be used to extract sources from the GSM and put them into a file that can directly be used as input for makesourcedb.

It can be run like:

makesourcedb key=value key=value ...

The possible keywords and values are described in the following table.

Keyword	Туре	Default	Description
in	string		Name of the input ASCII file
out	string		Name of the source table
outtype	string	casa	The output type of the source table which can be casa or blob. Casa is the standard type which can be used for all purposes. Blob is a special type that can only be used for NDPPP demixing. Its advantage is that it is much faster for the large source models used in demixing.
format	string	Name,Type,Ra,Dec,I,Q,U,V,SpInx, MajorAxis,MinorAxis,Orientation	The format of the ASCII file. It is described in detail in the next section.  A value " <filename" #(="" )="format&lt;/td" a="" be="" can="" file="" file.="" first="" form="" format="format-string" format-string="" forms;="" from="" given="" given,="" if="" in="" input="" is="" it="" line.="" means="" must="" name="" no="" preferred.="" read="" single="" specified="" string="" that="" the="" two=""></filename">
patchprefix	string	un	prefix for the patch name implicitly created from a source.
patchsuffix	string	un	suffix for the patch name implicitly created from a source.
append	bool	False	Append to the source table? If not, a new source table is created.
check	bool	False	Check for each entry if the patch or source name is duplicate? True makes the program somewhat slower. At the end it is always checked that no duplicate entries exist.

Keyword	Туре	Default	Description
center	angle		If given, only sources in the given circle or box around the center are written into the source table.  The center has to be given as "ra, dec" given in casacore's MVAngle format (discussed below).
radius	angle		Radius of a circle (in MVAngle format) around the center to select sources.
width	angle		Full width of a box around the center to select sources to be given as "width" (for a square box) or "rawidth, decwidth" (in MVAngle format).

makesourcedb reads lines from the input file, interprets them using the format string (described below), and writes the ones matching the optionally given circle or box into the Source Table. Possible comment lines (indicated by #), empty lines, and format line in the input file are ignored when reading the values from the file. Whitespace (one or more blanks and/or tabs) is ignored, unless used as separator. Leading and trailing whitespace is always ignored.

Adding a source to the source table means two things:

- it is added to the source table itself
- for all source parameters appropriate entries are added to the associated parm table. The names of those parameters look like sourcename: fieldname where field name is the name of a field given in the format string.

### Format string

The format string defines which fields are present in the ASCII file and which separators are used. For example:

#### Name, Ra, Dec

tells that the lines in the input file contain source name, ra, and dec. It also tells that the values on each line are separated by commas. A field name can consist of alphanumeric characters, underscores and colons. A colon cannot be the first character. All other characters (including whitespace) can be used as separator. A colon can also be used as field separator as long as it is surrounded by whitespace in the format string (to distinguish it from a colon in a field name).

A few special options are available in the format string.

• Default values are used for missing or empty field values in input lines. Standard default values (as listed in the table below) exist, but it is possible to specify an explicit default value in the format string. It can be done by giving a value (enclosed in quotes) for a field like:

• It is also possible to define a field with a fixed value that does not appear in the input lines like:

2025-09-09 01:50 3/7 makesourcedb

In this way the Category parameter will be defined for all sources, while the input lines do not contain the Category field.

• If the input lines contain fields that have to be ignored, they can be skipped by giving no field name in the format line or by using the field name dummy. For example:

means that the third field is skipped. Note that the name dummy has to be used if whitespace is used as separator like:

#### Name Ra dummy Dec

The format strings can be given in a few ways:

- Directly in the program format keyword.
- As a format line in the input file (before the first value line) or optionally in a separate text file. Such a format line has to start with format= (whitespace is allowed). For backward compatibility the format line can also be given as a special comment line like #(...)=format, where ... is the format string. Whitespace can be used.

### Field types

makesourcedb distinguishes two types of fields: standard fields and other fields. The standard fields have a predefined (case-insensitive) name. However, the name is case-sensitive in the sense that the name given in the table below is used as the parameter name in the parm table. The following standard fields are known:

Field	Туре	Unit	Default	Description
Name	string			Source Name
Туре	string			Source type. Currently only point, gaussian, or shapelet is possible.
Ra	angle			Right ascension in MVAngle format (see below)
Dec	angle			Declination in MVAngle format (see below)
I	double	Jy	1	Stokes I flux (for Gaussian sources this the <i>total</i> ' flux)
Q	double	Jy	0	Stokes Q flux
U	double	Jy	0	Stokes U flux
V	double	Jy	0	Stokes V flux
MajorAxis	double	arcsec	1	FWHM of major axis for a Gaussian source
MinorAxis	double	arcsec	1	FWHM of minor axis for a Gaussian source
Orientation	double	deg	1	Orientation of the major axis in degrees North over East
OrientationIsAbsolute	boolean		false	Whether the Orientation is absolute (w.r.t. declination axis) or relative (w.r.t. m axis in not-specified field). See the note below.

Field	Туре	Unit	Default	Description
IShapelet	string			Name of the ASCII file defining the shapelet coefficients and scale of the I flux in the 'Sarod' format. Note that field I gives the total flux, while IShapelet defines the shape.
QShapelet	string			
UShapelet	string		ldem for U flux.	
VShapelet	string		uu	Idem for V flux.
SpectralIndex	doubles		[]	Spectral index. Multiple values can be given enclosed in square brackets and separated by commas. The n-th value defines the n-th order term.  Note that the old way of specifying the spectral index (using SpectralIndexDegree and SpectralIndex:i) is not supported anymore. If used, an error is given.
ReferenceFrequency	double	Hz		Reference frequency (in Hz) for the spectral index
RotationMeasure	double	rad/m^2		Rotation measure for the polarized flux Q and U
PolarizationAngle	double	rad	0	Polarization angle (in radians) for the polarized flux
PolarizedFraction	double		0	Fraction of I that is polarized
ReferenceWavelength	double	m		Reference wavelength (in m) if RM is specified using Q and U
Category	integer		2	Source category (must be 1, 2, or 3)
Patch	string			Patch name
rah	double	hr	0	Special way to specify RA (see below)
rad	double	deg	0	
ram	double	min	0	
ras	double	sec	0	
dech	double	hr	0	Special way to specify DEC (see below)
decd	double	deg	0	
decm	double	min	0	
decs	double	sec	0	

All non-standard fields have type 'double' and a default of 0. The name as given in the format string is used as the parameter name in the parm table.

Note that the polarized flux can be given directly (as Q and U) or by means of the rotation measure, polarization angle, and polarized fraction. An error is given if both are given for the same source.

#### **Using Rotation Measure**

The flux of a source can be defined using a rotation measure by means of the field name RotationMeasure. The required polarization info can be given in two ways:

• Directly by means of the fields PolarizationAngle and PolarizedFraction. They define the angle and fraction for wavelength 0. If the field ReferenceWavelength is also given, it must 2025-09-09 01:50 5/7 makesourcedb

be 0.

• Indirectly by means of the Q and U flux. ReferenceWavelength should be given as well to define the wavelength they are valid for. If not given, it will be derived from the spectral index ReferenceFrequency. The polarization angle and fraction will be calculated taking the spectral index into account like:

```
polang = atan2(U,Q) / 2 - refwvl^2 * rm
polfrac = sqrt(Q^2 + U^2) / I(refwvl)
ReferenceWavelength must be > 0 if a spectral index is given.
```

### **Defining Patches**

A patch is defined when an input line contains a patch name and no source name. For a patch the fields Category, Ra, Dec, and I are also used which define its category, center, and total flux. In a future version of makesourcedb the center and total flux can be calculated from the sources that are part of the patch.

A source is attached to a patch if its input line contains a patch name. A line defining that patch should have been processed before this line.

If no patch name is given for a source (as is usually the case), the source forms a patch in itself and the source name is used as the patch name (with an optional prefix/suffix). In this way BBS can always deal with patches.

#### **Orientation of Gaussian sources**

Until 2022, the orientation of Gaussian sources has been interpreted somewhat wrong. It was interpreted (both by DP3 and pybdsf) as the angle with respect to the m axis of a certain image. The correct (implicit) definition is the orientation w.r.t. the declination axis. For fields far away from the NCP, and/or sources not too far from the field center, these two definitions are almost identical.

In DP3 v5.3 a fix was implemented to take the projection effects into account properly. This fix is only called when OrientationIsAbsolute is true. A script will be added to LSMTool to convert from 'relative' orientation to absolute ones (this script needs the phase center of the original image as an input).

## **Angle specification**

Angles can be given in a variety of ways using the MVAngle format available in casacore.

- In sexagesimal HMS format like 12:34:56.78 or 12h34m56.78. Note that a colon defines hours. To avoid mistakes, it is not allowed to use colons in declination values (but see below).
- In sexagesimal DMS format like 12.34.56.78 or 12d34m56.78.
- As a double value optionally followed by a unit like 12.34deg or 12.34rad. If no unit is given, radians are used.

In all cases negative values are possible. It is checked if RA is between -360 and +360 degrees and DEC between -90 and +90 degrees.

Some catalogues use a different way to specify Ra or Dec. For this reason the special field names rah, ram, and ras can be used in the format string. They define the hour, minute, and second part of

an RA. If the Ra is in degrees, rad should be used instead of rah. Similarly dech, decd, decm, and decs can be used for declination. A negative hour or degree part means that the entire value is negative. For example:

rah ram ras

can be used in the format string to interpret RA values like:

12 34 56.78 -12 34 56.78

Note that these special field names can also be used to interpret declination values given in degrees using colons. For instance:

decd : decm : decs

to interpret

12:34:56.78

# mergesourcedb

This program makes it possible to merge two SourceDB data sets or to copy (part of) a SourceDB. Patches for which the positions in both SourceDBs match (within a given uncertainty radius) are merged. It means that the patch in the second SourceDB will be used!! The mode parameter tells if non-matching patches in the second SourceDB should also be copied.

It is possible to select patches using a name pattern.

It can be run like:

mergesourcedb key=value key=value ...

The possible keywords and values are described in the following table.

Keyword	Туре	Default	Description		
in1	string		Name of the first SourceDB; this must be given		
in2	string	""	Name of the second SourceDB. If empty string, no merge will be done, but the patches of in1 are copied.		
out	string		Name of the output SourceDB		
outtype	string	Casa	The output type of the source table which can be casa or blob. Casa is the standard type which can be used for all purposes. Blob is a special type that can only be used for NDPPP demixing. Its advantage is that it is much faster for the large source models used in demixing.		
append	bool	False	Append to the source table? If not, a new source table is created.		
in1.patches	string	*	Pattern telling the names of the patches to take from in1		

Keyword	Туре	Default	Description
in2.patches	string	*	Pattern telling the names of the patches to take from in2
mode	string	all	match = only use in2 patches matching an in1 patch position and replace that in1 patch all = also copy the other in2 patches matching the pattern
radius	string	10arcsec	Uncertainty radius for matching patches in both SourceDBs

## showsourcedb

This program can be used to show the contents of a SourceDB data set.

It can be run like: It can be run like:

showsourcedb key=value key=value ...

The possible keywords and values are described in the following table.

Keyword	Туре	Default	Description
in	string		Name of the SourceDB; this must be given
mode	string	all	patch = show all patches source = show all sources all = show patches and sources skymodel = show all in skymodel format
patches	string	*	Pattern telling the names of the patches to take

From:

https://www.astron.nl/lofarwiki/ - LOFAR Wiki

Permanent link:

Last update: 2022-10-21 07:48

